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Correction methodology for the spectral interfering γ -rays overlapping to the analytical peaks used in the analysis of 232 Th

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ABSTRACT

In the γ -ray spectrometric analysis of the radionuclides, a correction factor is generally required for the spectral interfering γ -rays in determining the net areas of the analytical peaks because some interfering γ -rays often might contribute to the analytical peaks of interest. In present study, a correction methodology for the spectral interfering γ -rays (CSI) is described. In particular, in the analysis of 232 Th contained in samples, the interfering γ -rays due to 226 Ra, 235 U, 238 U and their decay products often overlap to the peaks of interest from 232 Th decay products, and vise versa. For the validation of the proposed CSI method, several certified reference materials (CRM) containing U and Th were measured by using a 76.5% efficient n-type Ge detector.

The required correction factors were quantified for spectral interference, self-absorption and true coincidence summing (TCS) effects for the relevant γ -rays. The measured results indicate that if one ignores the contributions of the interfering γ -rays to the analytical peaks at 583.2 keV of 208 TI and 727.3 keV of 212 Bi, this leads to a significantly systematic influence on the resulted activities of 232 Th. The correction factors required for spectral interference and TCS effects are estimated to be \sim 13.6% and \sim 15.4% for 583.2 keV peak. For the 727.3 keV peak, the correction factor is estimated to be \sim 15% for spectral interference, and \sim 5% for the TCS effects at the presently used detection geometry. On the other hand, the measured results also indicate that ignoring the contribution of the interfering γ -rays to the areas of the analytical peaks at 860.6 keV of 208 TI, 338.3 and 911.2 keV of 228 Ac does not lead to any significant systematic influence on the 232 Th analysis. Because these factors are remained generally less than \sim 5%, i.e., within overall uncertainty limits. The present study also showed that in view of both the spectral interference and TCS effects, the "cleaner" analytical peaks are seen at 338.3 keV (11.25%) and 911.2 keV (26.13%) of 228 Ac when high resolution γ -rays spectrometry was used in the 232 Th activity measurements. Therefore, they can be adopted as the "reference" peaks in the 232 Th analysis.

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1. Introduction

In practice, the high resolution γ -ray spectrometry is the most commonly used technique to determine activities or concentrations of radionuclides in a variety samples. Because it is quite easy, a less laborious and simple in view of sample preparation. Additionally it is a rapid and non-destructive method among all other nuclear analytical techniques. In this context, the γ -rays emitted from 228 Ac, 224 Ra, 212 Pb and 212 Bi nuclides in 232 Th decay series and 234 Th, 234 mpa, 226 Ra, 214 Pb, 214 Bi and 210 Pb nuclides in 238 U decay series, and 40 K nuclide can often be measured by using a suitable Ge detector. Although the modern γ -ray spectrometry is known as a mature technique to measure the radionuclides in a variety samples, it still needs to study carefully the activity correction factors affecting the resulted measurement accuracy.

This implies that the analytical procedures to be employed are carefully accounted for the radionuclides of interest to get more reliable measurement results. That is, if it is intended to perform the high-quality gamma spectrometric measurements, at least, the corrections for self-absorption effects for the measured γ -rays with energy of generally below 300 keV, true coincidence summing (TCS) effects for the multi-cascading γ -ray transitions, and the spectral gamma interferences to the analytical peaks must be taken into account in the used analytical procedures. Fortunately, there are numerous well-defined correction methods for the correction of TCS effects (Sudár, 2002; DSM, 2005; Arnold and Sima, 2006; Dryák and Kovár, 2009; Zhu et al., 2009; Yücel et al., 2009a) and self-absorption effects (ASTM E181, 2003; McMahon et al., 2004) from practical point of view. However, in literature survey, the methods relating to the estimation of the contributions from the interference peaks overlapping to the analytical peak of interest is still lacking or open to the investigation, which is important in view of obtaining more reliable peak area determinations. On the other hand, in recent years, a few

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complete papers have been published for the estimation of the spectral peak interferences to the analytical peaks used for ²³⁸U and ²²⁶Ra and their decay products (Papachristodoulou et al., 2003; De Corte et al., 2005; Yücel et al., 2009b; Yücel et al., 2010). Further, our previous investigations have also indicated that corrections for the spectral interferences (CSI) to an analytical peak needs to have more specific equations used for each peak but this can easily be implemented in the gamma spectrometric data evaluation by using the simple analytical procedures. However, it depends on several parameters such as γ -ray emission probabilities, detection efficiency, self-absorption and TCS effects. In the proposed CSI methodology, the complexity degree of the acquired γ -ray spectra is also very important for the estimation of the magnitudes of the spectral contributions to the analytical peak area of interest. It is clear that the use of Ge-detectors with the higher detector resolution (corresponding to smaller FWHM in keV) in the measurement setup allows to reduce remarkably the peak interferences to the analytical peak of interest.

On the other hand, ²³²Th activity is generally based on the quantification of the well-separated, a few prominent γ -ray peaks from its decay products. They are observed as the more intense γ rays at 238.6 keV (43.30%) of 212 Pb, 338.3 keV (11.25%) and 911.2 keV (26.13%) of 228 Ac, 727.3 keV (6.61%) of 212 Bi, 583.2 keV (30.52%) of 208 Tl and 860.6 keV (4.48%) of 208 Tl. There seems that these analytical peaks to be used in ²³²Th analysis might be suffered from the spectral interfering γ -ray emissions due to the presence of other nuclides from ²²⁶Ra, ²³⁵U and ²³⁸U besides ²³²Th series nuclides in all variety samples. At first sight, above six prominent peaks in ²³²Th decay series can be used reliably as analytical peaks in thorium analysis when γ -ray spectrometry used. However, it is worth noting that some of them still need to correct for especially the contribution of spectral interfering γ -rays and that of true coincidence-summing (TCS) effects. In the case of low-level activity measurements, for instance, in the measurement of environmental samples, it is fact that the close proximity between the source and detector is commonly employed to obtain the better counting statistics and/or shorten the measurement time, but this proximity also causes true coincidence summing effects (Sima and Arnold, 2000; Garciá-Talavera, 2003). The magnitudes of the TCS factors directly impact on the measurement accuracy in the γ -ray spectrometric measurements, due to decay scheme properties of some specific nuclides that have multi-cascading transitions, generally yielding to true summing effects. Hence the corrections for TCS effects must also be taken into account especially for the case of the close-in detection geometry conditions (Debertin and Helmer, 1988).

In view of the spectral interference correction, some authors (Papachristodoulou et al., 2003; De Corte et al., 2005) have previously noted that the least spectrally disturbed analytical peak among possible analytical peaks for ²³²Th analysis, i.e., almost an interference-free peak is seen at 338.3 keV of ²²⁸Ac. In other words, it may be adopted to be the "cleaner" peak for ²³²Th analysis from the point of view of the interfering γ -rays. Hence, this peak has already been chosen as a "reference peak" in our earlier study to correct for the spectral contribution of the 63.81 keV peak of ²³²Th to the 63.3 keV peak at of ²³⁴Th for the measurement of ²³⁸U activity (Yücel et al., 2009b). However, this "reference" analytical peak at 338.3 keV of 228Ac will be reexamined in detail for the purpose of ²³²Th analysis whether it has possible TCS effects in case of a close-in detection geometry and the influence on the net peak area due to its spectrally interfering γ -rays. Especially when the samples contained in much more U rather than Th, it can be questioned that whether the spectral contributions of ²²³Ra (²³⁵U) to the "reference" analytical peak at 338.3 keV of ²²⁸Ac (²³²Th) might be neglected or not. Additionally, it seems that self-absorption effects for above all six prominent analytical peaks used for the ²³²Th analysis cannot be considered a serious problem since they are lying in the intermediate energy region in the spectrum, however, their magnitudes cannot be omitted and should still be taken into account in a variety of the samples because these factors can sometimes have higher values depending on sample density and high Z-components in matrices.

Therefore, the main objective of the present study is to determine the contributions of spectral peak interferences to six prominent analytical peaks used in the direct measurement of $^{232}{\rm Th}$ activity in the samples. It is essentially aimed at proposing a useful method that will be used for the correction for the contribution of the spectral interfering γ -rays to an analytical peak. For the validation of a CSI methodology, the certified reference materials (CRM) containing a varying amount of U and Th were measured by using of a calibrated n-type Ge detector. From the measured γ -ray spectra of the CRM samples, the contribution of the interfering γ -ray emissions from decay products of $^{226}{\rm Ra}$, $^{235}{\rm U}$ and $^{238}{\rm U}$ overlapping to the analytical peak areas have been determined by using appropriate equations. The results obtained for the validation of CSI methodology are discussed in detail.

2. Experimental

The detector used was an n-type and coaxial high purity Ge (Ortec GMX70P4-S) with a 0.5 mm thick Be window. The detector has a measured relative efficiency of 76.5% and a peak-to-Compton ratio of 74:1 at 1332.5 keV (60 Co). The measured energy resolution is 2.08 keV at 1332.5 keV (60 Co) and 0.8 keV at 122.1 keV (57 Co). The Ge crystal has a 69.9 mm diameter and a 82.6 mm length, hole diameter of 9.2 mm and hole depth of 73.8 mm, the end-cap to crystal gap of 4 mm, the mount cup length of 105 mm, outside contact layer of 0.3 μ m boron, inner contact layer of 700 μ m lithium and insulator/shield of 0.05 mm aluminized mylar. The detector with a 1 mm thick Al end-cap was placed in a lead shield of 10 cm thickness, graded with 0.5 mm Sn and 1.6 mm Cu liners. The Pb shield is also jacketed by a 9.5 mm low-carbon steel outer casing. The base of lead shield has an annular lead plug in which accommodate only a dipstick cryostat and the detector cables.

The 1.5 mm thick plastic sample beakers with a 3.5 cm filling height and a 4.3 cm internal diameter were located onto the detector end-cap at a 2 mm distance by means of a sample locator. All powdered CRMs were individually filled in the plastic beakers and sealed tightly. Then they were kept for a period of more than one year. During the measurements, the nitrogen (N_2) gas-out boiling off from the dewar with a flow rate of about 0.25 l h⁻¹ through a flexible hose was flushed inside the shield in order to purge thoron 220 Rn (55.6 s) and radon 222 Rn (3.82 d) contamination around the sample.

The Ge detector was connected to a digital signal processing analyzer (ORTEC DSPEC jr. 2.0) with a 16 K ADC/MCA in which data acquisition and evaluation was performed through a gamma analysis software (Gamma Vision Ver. 6.01). The measurements for CRMs and standard were repeated at least three times. The sufficiently long periods varying between 24 h and 72 h were chosen for the CRM samples to obtain good statistics of the spectrum counts. A measured room background spectrum was always subtracted from the relevant peaks of CRM spectra. Dead times for the γ -ray spectra acquired for the CRMs were in the range 0.30–3.82%.

For the efficiency calibration of the present γ -ray spectrometer, a multi-nuclide standard source spiked in a sand (density:

The activities specified for the radionuclides in the certified reference materials.

Material description ^a	Code	Specific activity	Specific activity values certified for ra	ır radionuclides g,h (Bq kg $^{-1}$)	$\langle \mathbf{g}^{-1} \rangle$					
		²³⁰ Th	²²⁶ Ra	²¹⁰ Po	²¹⁰ pb	²³² Th	²²⁸ Ra	²²⁸ Th	²³⁸ U	²³⁵ U
Uranium tailing sample ^b Uranium tailing sample ^b Uranium-thorium ore ^c Uranium-thorium ore ^d Rare earth-thorium ore ^e Thorium ore ^f	UTS-1 UTS-2 DL-1a DH-1a OKA-2 RGTh-1	3600 ± 37 4400 ± 41 Not reported Not reported Not reported Not reported	3670 ± 38 5600 ± 46 1400 ± 40 31500 ± 1100 Not reported ⁱ Not reported ⁱ	3100 ± 35 4400 ± 41 Not reported Not reported Not reported Not reported	3250 ± 35 4550 ± 42 1400 ± 20 30800 ± 900 Not reported ⁱ Not reported ⁱ	680 ± 16 880 ± 18 308 ± 16 3692 ± 122 117381 ± 2353 3250 ± 90	680 ± 16 1000 ± 20 Not reported Not reported Not reported Not reported	710±17 920±19 Not reported ⁱ Not reported ⁱ Not reported ⁱ	Not reported ⁱ Not reported ⁱ 1432 ± 37 32462 ± 37 2699 ± 101 78 ± 6	Not reported ⁱ Not reported ⁱ 68 ± 2 1541 ± 2 128 ± 5 3.6 ± 0.3

b Recommended values for the elements or constituents in matrices are given within the ranges 2.62-4.87 wtt total Fe, 0.18-0.54 wtt Ti, 2.71-6.24 wtt Al, 0.42-5.24 wtt Ca, 0.23-1.23 wtt total S and 0.84-2.64 wtt a Certified Reference Materials (CRMs) were obtained from CANMET-Mining and Mineral Sciences Laboratories, Canada.

. The matrix given in its certificate is of a waste-grade material which is a pale yellow arkose sandstone having essentially the radioactive minerals.

d The ore matrix is described as a sericitic, feldspathic quartzite containing about 10% pyrite with major components of 79,75% SiO₂, 5.17% Fe, 4.8.2% S and 3.44% Al in its certificate

e Britholite ore contains 3.292% wt ThO2 and its remaining components consist of mainly 33.51% rare earth (RE) elements oxides (RE₂O₃+Y₂O₃), 14.72% SiO₂, 25.32% CaO and 5.83% Fe₂O₃.

This is 1AEA reference material, prepared from BL-5 (7.9% U) and OKA-2 (2.89% Th, 219 µg U/g) CRMs by floating silica powder of similar grain size distribution. g All specific activities are given in a 95% confidence level.

h The activity conversions are 232Th=4057.4 ± 17 Bq in 1 g natural thorium, 238U=12,347.5 ± 6.7 Bq, 234U=12,223.4 ± 4.5 Bq and 235U=586 ± 1 Bq in 1 g natural uranium in which the data are based on IUPAC and NNDC

i No concentration or activity values were quoted in the certificates.

 $1.7\pm0.1\,\mathrm{g\,cm^{-3}}$ of SiO₂) matrix containing $^{210}\text{Pb},~^{109}\text{Cd},~^{57}\text{Co},$ $^{123m}\text{Te},~^{51}\text{Cr},~^{113}\text{Sn},~^{85}\text{Sr},~^{137}\text{Cs},~^{88}\text{Y},~^{60}\text{Co}$ nuclides was used to obtain the full-energy peak (FEP) efficiency curve. The standard source used was purchased from Isotope Products Inc., traceable to NIST. The specific activities for the radionuclides contained in the CRMs are given in Table 1.

3. Correction methodology for spectral interferences

The CSI methodology is implemented for the correction for the contributions of spectrally interfering γ -ray emission(s), i.e., those overlapping to an analytical peak of interest. If the interfering γ -rays are close to the centroid energy of the measured sum peak (or the mixed peak) within the width of $1 \times FWHM$ (full width at half maximum), they are assumed to be overlapped to the analytical peak of interest. FWHM is equal to $2.355 \times \sigma$, where σ is a width parameter (standard deviation of the Gaussian distribution representing a full-energy peak). It is important to note that when this proximity between two peaks is greater than $1 \times FWHM$ to the peak centroid energy of the interested peak, the CSI approach does not involve those distant peaks, which are outside the limits of a full-energy peak width of $1 \times FWHM$. They are not deemed the spectrally interfering peaks overlapping to the peak of interest. Because, in practice, when the sensitivity test in a spectrum region is passed the factor of between 2 and 3.3 times FWHM width, the spectrum peak region is often de-convoluted to obtain, at least, two fitted peaks lying in that region by use of the gamma analysis software. Therefore, the interfering γ -rays lying in the width $(1 \times FWHM)$ of the mixed peak are taken into account in present CSI methodology. Assuming that the peak width is 1 keV in low energy region of below 300 keV and 1.5 keV for the intermediate energy region of 300-1000 keV and 2 keV in FWHM (in keV) for the high energy region of 1000-2000 keV in the γ -ray spectrum. The chosen energy regions in the spectrum and the peak widths of interest can be altered arbitrarily, depending on mainly the energy resolution of the used detector. Hence it is clear that the proposed CSI method will always be valid for the standard γ -ray spectrometry equipped with the commonly used Ge-detectors with an energy resolution of around 2 keV at 1332.5 keV of 60Co. In fact, the present CSI method is basically a peak stripping method (but not the "spectrum stripping") as a means of avoiding de-convolution. Thus it is worth noting that the uncertainties in de-convolution process and the difficulty of checking the performance whether it can be avoided altogether. Therefore, CSI methodology can be applied reliably to the "singlet" peaks. However, it is possible that CSI method can also be applied to any "multiplet" on the provided that the considered "multiplet" peak be deconvoluted by using a suitable peak fitting program. An example of the two close peaks is shown in Fig. 1, in which 238.6 keV peak of ²¹²Pb (²³²Th) is deconvoluted successfully from its nearest peak at 242 keV of ²¹⁴Pb(²²⁶Ra) by peak fitting engine of the used gamma software. In this case, the sensitivity test of the spectrum analysis program requires making the peak de-convolution because the proximity between two nearest peaks is lying in $\sim 2.3 \times FWHM$ in the present peak region, that is greater than 2 × FWHM. Hence, after that deconvolution process, the user should be grouped the interfering γ rays in two separate "fitted" peaks and their interfering peaks must be, in turn, evaluated from the fitted peak area of 238.6 keV peak and that of 242 keV peak by employing CSI methodology.

The proposed CSI is also a simple calculation method based on the net peak areas determined from one measured γ -ray spectrum to obtain more accurate the net peak area of the interested analytical peak. For the figurative illustration of the CSI method, a schematic representation of a mixed peak is shown in Fig. 2

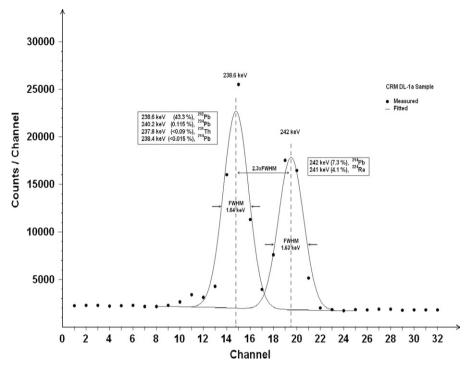


Fig. 1. Deconvolution of 238.6 keV peak of ²¹²Pb from its nearest peak at 242.0 keV peak of ²¹⁴Pb observed in the spectrum region of CRM DL-1a sample

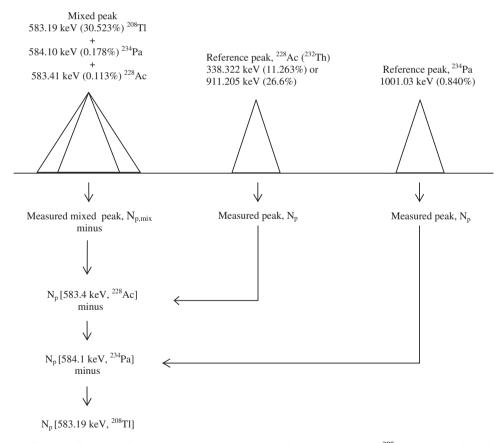


Fig. 2. The figurative illustration of the peak area correction method on the 583.19 keV peak of ²⁰⁸Tl for its spectral interferences.

for the 583.2 keV peak of 208 Tl in which there are two interfering γ -ray emissions contributing to that analytical peak.

Then, the magnitude of the spectral interference degree (or the percentage contribution) of any overlapping γ -rays to the measured mixed peak can be defined as the ratio of the net peak areas,

i.e., its proportion in the measured mixed peak, r (%)

$$r = 100 \times \left[\frac{N_{p,i}(E_i)}{N_{p,mixed}(E)} \right]$$
 (1)

where $N_{p,i}(E_i)$ is the net peak area of the i-th peak (p) with energy E_i from a given nuclide and $N_{p,mixed}(E)$ is the net peak area of the mixed peak with the centroid energy E for the all emissions lying in the measured full-energy peak region of interest. The sum of the peak areas of all γ -ray emissions overlapping to the mixed peak is

$$N_{p,mixed}(E) = \sum_{i=1}^{n} N_{p,i}(E_i)$$
(2)

Therefore, in the proposed CSI methodology, the first step is to review and determine the interfering γ -rays and their emission probabilities from the relevant decay databases. The second step is to determine accurately the net peak counts for each interfering γ -ray by choosing a suitable "reference" peak. Then, third step is to establish a suitable CSI equation(s) for the determination of the contribution of any interfering γ -ray emission. Such an interference correction equation for a particular interfering γ -ray emission can be written in the following form:

$$N_{p,i}[E_i,nuclide]_{inf} = \frac{[I_{\gamma}\varepsilon_p F_c F_s]_{E_i,inf}}{[I_{\gamma}\varepsilon_p F_c F_s]_{E,ref}} \times N_{p,ref}[E,nuclide]_{ref}$$
(3)

where $N_{p,i}$ is the net peak area in which the subscript "inf" represents the interfering nuclide and $N_{p,ref}$ is the net peak area of the reference peak in which the subscript "ref" represents a particular peak chosen as a "reference" peak. ε_p is FEP efficiency, I_γ is γ -ray emission probability, F_s is self-absorption correction factor and F_c is TCS correction factor at a given energy E (or E_i). In each equation, the particular "reference" peak(s) as given in Table 2 is chosen deliberately from the spectrally undisturbed peaks among all possible peaks that can be used for 226 Ra, 235 U, 238 U and 232 Th and their decay products used in analysis.

The "reference" peak should also be chosen from the fairly intense and TCS-free $\gamma\text{-rays}$ when the relevant decay series is examined

carefully. However, it is important to note that when the "reference" peak is chosen from the peaks of 235 U nuclide, the right side of Eq. (3) must be divided by a factor, f corresponding to the activity ratio of 238 U to 235 U, $f=(A^{238}/A^{235})=([\theta\times\lambda]_{238}/[\theta\times\lambda]_{235})\times(M^{235}/M^{238}),$ in which A is the activity, θ is the isotopic abundance, λ is the decay constant, M is the atomic mass and "235" and "238" refer to 235 U and 238 U, respectively. In the case of natural uranium $f=21.72\pm0.03$ is taken. It is assumed that 226 Ra is in equilibrium with 238 U and the activity ratio of $f=A^{238}/A^{235}$ to be natural in samples. Otherwise, the misleading results might be obtained for the net peak area, N_p [185.7 keV, 235 U].

For 232 Th analysis, Eq. (3) is arranged suitably for the determination peak areas $N_{p,i}(E_i)$ of the spectral interfering peaks overlapping to the measured analytical peak areas, $N_{p,mixed}(E)$ used for 232 Th analysis by using appropriate reference peaks $N_{p,ref}$ as given in Table 2, whichever appropriate. Then these contributions $N_{p,i}(E_i)$ are calculated for each of the interfering peaks and expressed the peak proportion of the mixed peak area $N_{p,mixed}(E)$ to obtain the final more accurate net area for the analytical peak of interest. However, when the γ -ray peaks from NORM nuclides are to be used in the analysis it is imperative that the peaked-background area be subtracted from the net peak area for a given measured peak, as noted by Gilmore (2008).

4. Results and discussion

The γ -ray spectrometer with an n-type Ge detector was calibrated by measuring the multi-nuclide standard source at a given geometry to obtain a FEP efficiency curve in the energy range from 46.5 to 1836 keV. In the present efficiency calibration, the TCS effects on the apparent ε_p efficiency values for particular peaks of 60 Co and 88 Y was calculated by use of a GESPECOR® (Ver 4.2)

Table 2The reference peaks used in Eq. (3) for the correction of the interfering peaks to the analytical peaks of interest.

Analytical peak of interest $N_{p,mixed}$ (E)	Interfering peak area $N_{p,i}(E_i)$	Reference peak area $^{\circ}$ $N_{p,ref}$ (E)	Note
238.6 keV (43.3%), ²¹² Pb	237.8 keV (< 0.09%), ²³¹ Th	185.7 keV (57.1%), ²³⁵ U	The right side of Eq. (3) is divided by the factor, <i>f</i>
	238.4 keV (< 0.015%), ²¹⁴ Pb	295.2 keV (18.412%) or 351.9 keV (35.59%), ²¹⁴ Pb	-
	240.2 keV (0.515%), ²³⁴ Pa	1001.0 keV (0.84%), ^{234m} Pa	=
241.0 keV (4.1%), ²²⁴ Ra	242. keV (7.3%), ²¹⁴ Pb	295.2 keV (18.412%) or 351.9 keV (35.59%), ²¹⁴ Pb	In the ²³² Th analysis, 241.0 keV peak is not preferred as an analytical peak due to its weak intense and having a serious spectral interfering problem.
338.3 keV (11.263%), ²²⁸ Ac	338.28 keV (2.795%), ²²³ Ra	185.7 keV (57.1%), ²³⁵ U	The right side of Eq. (3) is divided by the factor, f
583.2 keV (30.523%), ²⁰⁸ Tl	583.41 keV (0.113%), ²²⁸ Ac	338.3 keV (11.263%) or 911.2 keV (26.6%), ²²⁸ Ac	The branching ratio of 36% for ²⁰⁸ Tl (²³² Th) is considered
	584.10 keV (0.178%), ²³⁴ Pa	1001.0 keV (0.84%), ^{234m} Pa	<u>-</u>
727.3 keV (6.605%), ²¹² Bi	726.86 keV (0.629%), ²²⁸ Ac	338.3 keV (11.263%) or 911.2 keV (26.6%), ²²⁸ Ac	-
	727.80 keV (0.050%), ²¹⁴ Bi	295.2 keV (18.412%) or 351.9 keV (35.59%), ²¹⁴ Pb	-
	727.80 keV (0.113%), ²³⁴ Pa	1001.0 keV (0.84%), ^{234m} Pa	-
860.6 keV (4.484%), ²⁰⁸ Tl	860.0 keV(0.001%), ²¹⁰ Tl	295.2 keV (18.412%) or 351.9 keV (35.59%), ²¹⁴ Pb	The branching ratio of 36% for 208 Tl (232 Th) is considered, but that of $2\times 10^{-2}\%$ for 210 Tl (238 U) is negligibly small
911.2 keV (26.6%), ²²⁸ Ac	910 keV (2.969%), ²¹⁰ Tl	295.2 keV (18.412%) or 351.9 keV (35.59%), ²¹⁴ Pb	The branching ratio of $2 \times 10^{-2}\%$ for ^{210}Tl (^{238}U) is negligibly small

^a Interfering Peak area $N_{p,i}\left(E_{i}\right)$ is calculated from Eq. (3).

^b The net area of the reference peak $N_{p,ref}(E)$ will be used in Eq. (3).

program in which a Monte Carlo modeling of the present detectorsample geometry was performed, taking into account the surrounding shield materials. This program was also applied to obtain the self-absorption F_s and TCS effects F_c correction factors for the relevant γ -rays for both the standard source and all CRMs used in the present analysis. Since the exact compositions of the measured CRMs and the standard source are known, the self-absorption factors F_s were also calculated for the used cylinder beaker by the well known approximation (ASTM E181, 2003) using a simple relation $F_s = \mu x/[1-\exp(-\mu x)]$ where μ (cm⁻¹) is linear attenuation coefficient and x (cm) is effective sample thickness $x = m/(\rho S)$ in which ρ (g cm⁻³) is sample density. S is the extended-source area. Mass attenuation coefficients (u/ρ) are taken from both KORDATEN database, which is used in GESPECOR® program, and the wellknown NIST-XCOM database (XCOM, 2009). The self-absorption F_s factors calculated by using two different databases were then averaged. The resulted factors for the self-absorption F_s and TCS effects F_c for all relevant γ -rays are given in Table 3. The estimated TCS effects are relatively smaller on the measurable analytical peaks at 238.6 keV of 212 Pb and 338.3 and 911.2 keV of 228 Ac but those F_c factors are relatively larger for 583.2 keV of ²⁰⁸Tl and 727.3 keV of ²¹²Bi due to their decay scheme properties. Since the present measurements were performed in a close-in detection geometry,

their magnitudes are important in the resulted activities. However, it is important to note that even if both the factors F_c and F_s are smaller, they cannot be neglected in CSI methodology because they are also involved in the CSI equations.

The uncertainty sources in the experimental data are mainly due to the counting statistics (0.03–5.5%), the detection efficiency (2.5–3.6%), weight determinations (0.01%), self-absorption and TCS factors (1.7–2.6%) and the uncertainties (0.03–7.7%) in the certified activities. They were combined using the usual law of uncertainty propagation according to ISO (1995) and EURACEM/CTAC Guides (2000). There are the additional systematic uncertainties due to the γ -ray emission probabilities (0.30–1.2%) and the variations in sample heights and homogeneity (2.1%) that were included in the overall combined uncertainty. The standard uncertainty of the combined result from three independent measurements is calculated from the pooled (internal) variance.

The measured activities were also averaged from three independent measurements for each CRM. The activity results without and with corrected activity for spectral interferences to the analytical peaks are given in Table 4 for ²³²Th analysis. The corrected activities for spectral interferences to each analytical peak are compared with those certified activities for ²³²Th in some CRMs. The uncertainties for the activity results given

Table 3 True coincidence summing and self-absorption factors for the γ -rays used in the ²³²Th analysis.

Nuclide	Energy E (keV)	γ -Ray emission probability, I_{γ} (%)	True coincidence summing factor ^a , F_c	Self-abso	Self-absorption factor $^{\rm b}$, F_s							
				UTS-1	UTS-2	DL-1a	DH-1a	OKA-2	RGTh-1 ^c			
²¹⁴ Pb	238.4	0.015	0.825	1.34	1.32	1.24	1.26	1.24	1.24			
²¹² Pb	238.63 ^d	43.3	0.999	1.34	1.32	1.24	1.26	1.24	1.24			
²³⁴ Pa	240.20	0.515	0.553	1.34	1.32	1.24	1.25	1.24	1.24			
²²³ Ra	338.28	2.795	0.999	1.29	1.27	1.21	1.22	1.20	1.21			
²²⁸ Ac	338.32 ^d	11.252	0.959	1.29	1.27	1.21	1.22	1.20	1.21			
²⁰⁸ Tl	583.19 ^d	84.952	0.867	1.22	1.21	1.17	1.17	1.16	1.16			
²²⁸ Ac	583.41	0,113	0.710	1.22	1.21	1.17	1.17	1.16	1.16			
²³⁴ Pa	584.1	0.175	0.830	1.22	1.21	1.17	1.17	1.16	1.15			
²²⁸ Ac	726.86	0.629	0.809	1.20	1.19	1.15	1.16	1.14	1.15			
²¹² Bi	727.33 ^d	6.605	0.957	1.20	1.19	1.15	1.16	1.14	1.15			
²³⁴ Pa	727.8	0.113	0.800	1.20	1.19	1.15	1.16	1.14	1.15			
²¹⁴ Bi	727.8	0.050	_	1.20	1.19	1.15	1.16	1.14	1.15			
²¹⁰ Tl	860	0.001	_	1.18	1.17	1.14	1.15	1.13	1.14			
²⁰⁸ Tl	860.56 ^d	4.484	0.968	1.18	1.17	1.14	1.15	1.13	1.14			
²²⁸ Ac	911.20 ^d	26.13	0.967	1.18	1.17	1.13	1.14	1.12	1.13			
²¹⁰ Tl	910	2.969	_	1.18	1.17	1.13	1.14	1.12	1.13			

^a True coincidence-summing factors were calculated for a right cylinder sample placed on the end-cap of an n-type, 76.5% relative efficient Ge detector.

Table 4 The measured activity results for 232 Th in CRMs without and with spectral interference correction.

CRM Certified activity ^a (Bq kg ⁻¹)		DL-1a (308 ± 16)			DH-1a (3692 ± 122))		RGTh-1 (3250 ± 90)		
Nuclide	Energy (keV)	Measured activity (Bq kg ⁻¹)	Corrected activity (Bq kg ⁻¹)	%∆ ^b	Measured activity (Bq kg ⁻¹)	Corrected activity (Bq kg ⁻¹)	%∆ ^b	Measured activity (Bq kg ⁻¹)	Corrected activity (Bq kg ⁻¹)	%Д ^b
²¹² Pb	238.6	341 ± 17	338 ± 17	-9.7	2076 ± 104	4045 ± 104	-9.6	3494 ± 196	3142 ± 180	-3.3
²²⁸ Ac	338.3	328 ± 18	325 ± 18	-5.5	3897 ± 200	3725 ± 200	-0.9	3408 ± 195	3176 ± 188	-2.3
²⁰⁸ Tl	583.2	296 ± 15	288 ± 15	6.5	4007 ± 201	3939 ± 200	-6.7	3047 ± 171	2987 ± 169	-8.1
²¹² Bi	727.3	319 ± 19	276 ± 19	10.4	3898 ± 198	3161 ± 198	14.4	3431 ± 194	2967 ± 193	-8.7
²⁰⁸ Tl	860.5	282 ± 21	282 ± 21	8.4	3478 ± 179	3478 ± 179	5.8	3099 ± 175	3086 ± 174	-5.1
²²⁸ Ac	911.2	303 ± 16	303 ± 16	-1.6	3675 ± 189	3675 ± 189	0.5	3188 ± 182	3176 ± 182	-2.3

 $^{^{\}rm a}$ Uncertainties are given within $\,\pm\,2\sigma$ in this table.

^b Self-absorption factors were calculated for CRMs using both GESPECOR® program and the approach given in ASTM E181 (2003).

^c The matrix for IAEA RGTh-1 reference standard material was assumed to be 99% silica for self-absorption factors. ^d The analytical peak of interest used in the present ²³²Th analysis.

 $^{^{\}rm b}$ % \varDelta is the percentage difference between the corrected activity and the certified activity for 232 Th given in Table 1.

in Table 4 are reported within $\pm 2\sigma$ limits (95% confidence level) to compare with the uncertainties in their respective certified activities of CRMs.

The resulted contributions were calculated from Eq. (3) for the spectrally interfering peaks and they are given in Table 5 as the peak area proportion of the measured (mixed) peak. Since peak area proportion is defined as the measure of spectral contribution for a peak by Eq. (2), and the measured range for peak area proportion are obtained from the replicate γ -ray spectra taken for UTS-1, UTS-2, DL-1a, DH-1a, RGTh-1 and OKA-2. The resulted contribution of the spectral interfering peaks are also indicated in Table 5 as the measured range of peak area proportion. The reason for this range is that the measured peak areas and the calculated interfering peak areas varied slightly from one CRM sample to other type CRM sample although they are measured in a fixed counting conditions. This is mainly due the complexity of the spectrum of each CRM, which might depend on several parameters. The spectrum shape of the measured peak are somehow affected by the radionuclide contents contained in samples, for instance, due to the amounts of U and Ra in addition to Th in samples. In other words, the absence or presence of an interfering radionuclide or its activity level besides the measured radionuclide of interest affects the spectrum complexity. Further, the activity degree of each radionuclide in sample plays an important role to this contribution (Yücel et al., 2009b). Hence the results for peak area proportion in Table 5 are weighted for each peak from those peak proportions, r in the range of contributions. The analyst might take into account these ranges of "interference" should be sure about the range of the real activity of the respective radionuclides in the sample. This finding also implies that the percentage peak area contribution of the interfering peak to an analytical peak is not a universal constant for the peak but it also

depends on the sample self-absorption and TCS effects, thus the latter two factors affecting the peak "interference" influence. Hence, the correction for the interfering peaks to each analytical peak must be performed for each measured sample by employing CSI methodology. It is worth noting that in case the composition of the used sample is unknown, the relevant F_s factors for the γ -ray energies should be determined from the well-known practical procedures in the literature. The TCS factors of the used γ -ray energies are also readily determined at a given counting geometry by employing one of several TCS correction methods (Sudár, 2002; Dryák and Kovár, 2009; Yücel et al., 2009a; Zhu et al., 2009).

The present results have also indicated that ignoring of the contribution of the interfering γ -rays to the peak areas of the analytical peaks at 238.6 keV of ²¹²Pb, 860.6 keV of ²⁰⁸Tl, 338.3 and 911.2 keV of ²²⁸Ac does not lead to any significant systematic influence on the ²³²Th activity analysis. Because the factors obtained for their spectral interference γ -rays are remained generally less than \sim 5%, i.e., within overall estimated uncertainty limits. But there is an important point for the analysis of 238.6 keV peak of ²¹²Pb (²³²Th), which is very close to the composite peak at 241 keV(4.1%) of ²²⁴Ra plus 242 keV (7.3%) of ²¹⁴Pb, as shown in Fig. 1. Depending mainly on the energy resolution of the detector, it should be deconvoluted appropriately by a peak fitting program before the CSI method is applied to its spectral interference correction. Otherwise, this close peak often might suffer the 238.6 keV peak of ²¹²Pb (²³²Th), especially when higher ^{224,226}Ra contents contained in the samples.

On the other hand, if one ignores the contributions of the interfering γ -rays to the analytical peaks at 583.2 keV of 208 Tl and 727.3 keV of 212 Bi, this leads to a much significantly systematic influence on the measured activities of 232 Th due to spectral

Table 5Peak area proportions for the peaks used in ²³²Th analysis determined by CSI methodology.

Nuclide	Energy (keV)	γ -ray emission probability I_{γ} (%)	Measured range of peak area proportion ^a (%)	Mean value for peak area proportion ^b (%)	Note
²¹² Pb(²³² Th)	238.63 ^f	43.3(8)	97.85-99.99	99.06	No correction is required for TCS effect but required a spectral interference correction after it was deconvoluted from the nearest peak at 241 keV (4.1%) of ²²⁴ Ra+242 keV (7.3%) of ²¹⁴ Pb, depending on the detector resolution
234 Pa(238 U)	240.20	0.515(2)	0.01-1.71	0.65	-
214 Pb(226 Ra)	238.4	< 0.015	0.01-0.44	0.29	-
²²⁸ Ac(²³² Th)	338.32 ^f	11.252(10)	95.77-99.99	99.11	Correction required slightly for TCS effect but no required for spectral interference
223 Ra(235 U)	338.28	2.795(4)	0.01-4.23	0.89	-
²⁰⁸ Tl(²³² Th)	583.19 ^f	84.952(7) ^c	98.29-99.69	99.12	Corrections required for both TCS effect and spectral interference
²³⁴ Pa(²³⁵ U)	584.1	0.175(2)	0.01-1.44	0.59	_
²²⁸ Ac(²³² Th)	583.41	0.113(4)	0.27-0.31	0.29	_
²¹² Bi(²³² Th)	727.33 ^f	6.605(8)	76.54-93.05	85.15	Corrections required for both TCS effect and spectral interference
²²⁸ Ac(²³² Th)	726.86	0.629(3)	6.85-7.54	7.18	-
²³⁴ Pa(²³⁸ U)	727.8	0.113(1)	1.33-5.82	3.90	-
²¹⁴ Bi(²³⁸ U)	727.8	0.050	0.01-5.62	3.77	_
²⁰⁸ Tl(²³² Th)	860.564 ^f	12.424(1) ^c	100	100	Correction required for TCS effect but no required for spectral interference
²¹⁰ Tl(²³⁸ U)	860	$6.90(20)^{d}$	e	e	_
²²⁸ Ac(²³² Th)	911.205 ^f	26.130(4)	100	100	Correction required slightly for TCS effect but no required for spectral interference
²¹⁰ Tl(²³⁸ U)	910	2.969(2) ^d	e	e	-

^a Peak area proportion is the measure of spectral contribution for the peak as defined by Eq. (2), and the measured range are obtained from the replicate gamma-ray spectra of RGTh-1, DL-1a, DH-1a, UTS-1, UTS-2 and OKA-2.

b The percentage mean values are averaged from the peak area proportions obtained from RGTh-1, DL-1a, DH-1a, UTS-1, UTS-2 and OKA-2.

^c A branching ratio of 35.94% is applied to estimate ²³²Th activity.

^d A branching ratio of 0.02% is applied to estimate ²³⁸U activity.

e Since the branching ratio of 0.02% is very low for ²¹⁰Tl, the peaks at 860 and 910 keV of ²¹⁰Tl did not contribute to the relevant analytical peaks.

f The analytical peaks of interest.

interferences. Thus, the correction factors are estimated to be $\sim\!15.4\%$ for TCS effects from Table 3 and $\sim\!13.6\%$ for spectral interferences from Table 5 for 583.2 keV peak. For the case of 727.3 keV peak, the correction factors are $\sim\!5\%$ for TCS effects and $\sim\!15\%$ for spectral interferences. Additionally, the self-absorption factors F_s have to be taken into account in the present γ -ray measurements since the variety of matrices and densities of CRMs are different from that of the multi-nuclide standard used in this work.

The present results also indicate that, in view of both the spectral interference and TCS effects the "cleaner" analytical peaks are measured at 338.3 and 911.2 keV of ²²⁸Ac. It is found that these two peaks can be adopted as the "reference" peaks in all spectral corrections, and even they can reliably be suggested as "analytical peaks" for ²³²Th activity analysis in almost all measurement conditions.

5. Conclusions

The present results indicate that if one ignores the spectral interference contributions from ²²⁶Ra, ²³⁵U and ²³⁸U nuclides to the ²³²Th nuclides activity, the systematic influences arising from their spectral interfering γ -rays to an analytical peak jeopardizes to obtain the true activity in sample. This can, however, be eliminated by using the presently proposed CSI method. Because it incorporates both the self-absorption and TCS correction factors in addition to the γ -ray emission probabilities. In CSI method, it is imperative that the particular equations be used for each analytical peak. However, it is evident that the emission probabilities of the interfering γ -rays to the measured (mixed) peak affect mainly the magnitude of spectral interference correction. In this work, the CSI methodology has been formulated only for the determination of contributions from ²²⁶Ra, ²³⁵U, ²³⁸U-decay series to ²³²Th decay products in the samples. It is worth noting that the magnitude of spectral interference contribution cannot assumed to be a universal, fixed value for a nuclide because it depends on not only the complexity of the acquired spectrum and the energy resolution of the detection system but also the radionuclide contents in the measured sample.

The proposed CSI methodology might be used as a post-it analysis tool for obtaining high-quality γ -ray spectrometric measurements of radionuclides contained in almost every type of the samples.

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